

Krylov Complexity in Free and Interacting Scalar Quantum Field Theories

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• Motivation: Krylov complexity is a measure of operator growth which has been shown to capture signatures of chaos in quantum many-body systems [Parker, et al. (2018)].

• <u>Abstract:</u> In <u>arXiv:2212.14702</u> (with Viktor Jahnke, Keun-Young Kim and Mitsuhiro Nishida), we study Krylov complexity in free and interacting massive scalar QFTs in $\mathbb{R}^{1,d-1}$ at finite temperature.

(1) Lanczos Coefficients: a_n, b_n

(2) Krylov complexity $K_{\hat{O}}(t)$ and <u>quantum chaos</u>

(3) <u>Krylov complexity</u> $K_{\hat{\alpha}}(t)$ in QFTs

Q: What is the behavior of b_n and Krylov complexity in QFTs?

We are interested in studying the time evolution of Heisenberg operators $\mathcal{O}(t)$ in a quantum system with Hamiltonian \hat{H} :

$$\begin{split} \hat{\mathcal{O}}(t) &= e^{it\hat{H}}\hat{\mathcal{O}}e^{-it\hat{H}} = \hat{\mathcal{O}} + it[\hat{H},\hat{\mathcal{O}}] + \frac{(it)^2}{2!}[\hat{H},[\hat{H},\hat{\mathcal{O}}]] + \cdots \\ &= \sum_{n=0}^{\infty} \frac{(it)^n}{n!} \mathscr{L}^n \hat{\mathcal{O}} \quad , \quad \text{where} \quad \mathscr{L} := [\hat{H},\cdot] \text{ is called the} \end{split}$$

Liouvillian.

With \mathscr{L} , we can define a Gelfand-Naimark-Segal (GNS) Hilbert space after choosing an inner product $(\mathscr{L}^m \mathcal{O} | \mathscr{L}^n \mathcal{O})$.

In general, the basis $|\mathscr{L}^n \hat{\mathcal{O}}\rangle$ is not orthonormal. However, we can find an orthonormal basis, known as the Krylov **basis** $(\hat{\mathcal{O}}_m | \hat{\mathcal{O}}_n) = \delta_{mn}$ using the Gram-Schmidt procedure:

 $|\hat{\mathcal{O}}_{-1}\rangle := |\mathbf{0}\rangle$, $|\hat{\mathcal{O}}_{0}\rangle := |\hat{\mathcal{O}}\rangle$

$$|\hat{\mathcal{O}}_{n}\rangle = \frac{1}{b_{n}} \left((\mathscr{L} - a_{n-1}) |\hat{\mathcal{O}}_{n-1}\rangle - b_{n-1} |\hat{\mathcal{O}}_{n-2}\rangle \right) \quad (n \ge 1)$$

In the Krylov basis, the matrix elements of the Liouvillian take the form

$$(\hat{\mathcal{O}}_m | \mathscr{L} | \hat{\mathcal{O}}_n) = \begin{pmatrix} a_0 & b_1 & 0 & 0 & \cdots \\ b_1 & a_1 & b_2 & 0 & \cdots \\ 0 & b_2 & a_2 & b_3 & \cdots \\ 0 & 0 & b_3 & a_3 & \ddots \\ \vdots & \vdots & \vdots & \ddots & \ddots \end{pmatrix}$$

In the Krylov basis, the time-evolved operator can be written as $|\hat{\mathcal{O}}(t)\rangle = \sum_{n=1}^{\infty} i^n \varphi_n(t) |\hat{\mathcal{O}}_n\rangle$, where: $\varphi_n(t) = \frac{1}{b_n} \left(ia_{n-1}\varphi_{n-1}(t) - \partial_t \varphi_{n-1}(t) + b_{n-1}\varphi_{n-2}(t) \right) \quad (n \ge 1)$ The *Krylov complexity* of the operator \hat{O} is defined as

$$K_{\hat{\mathcal{O}}}(t) := \sum_{n=0}^{\infty} n |\varphi_n(t)|^2$$

Krylov complexity measures the growth or "spread" of the operator \mathcal{O} in *Krylov (sub-)space* $\mathcal{H}_{\hat{\mathcal{O}}}$, the space spanned by the Krylov basis $\{ | \mathcal{O}_n \} \}$.

Chaos in Quantum Many-Body Systems

In [Parker, et al. (2018)] it was conjectured that in the thermodynamic limit, the b_n of generic non-integrable quantum many-body systems with local interactions should grow as fast as possible,

 $b_n \sim \alpha n + \gamma \quad (n \to \infty)$

In this case, the Krylov complexity should have an exponential growth $K_{\hat{O}}(t) \sim e^{\lambda_K t}$, with growth-rate λ_K To answer this, in <u>arXiv:2212.14702</u> we focus on the Wightman-ordered two-point function for a scalar field:

 $\Pi^{W}(t, \mathbf{x}) := \langle \phi(t - i\beta/2, \mathbf{x})\phi(0, \mathbf{0}) \rangle_{\beta}$

where $\langle \cdot \rangle_{\beta}$ is the thermal expectation value. Its Fourier transform is the *Wightman power spectrum*

$$f^{W}(\omega) := \int dt \,\Pi^{W}(t, \mathbf{0}) e^{i\omega t} = \frac{1}{\sinh(\beta\omega/2)} \int \frac{d^{d-1}\mathbf{k}}{(2\pi)^{d-1}} \rho(\omega, \mathbf{k})$$

where $\rho(\omega, \mathbf{k})$ is the spectral density. From the power spectrum, it is possible to compute the *moments*

$$u_{2n} = \frac{1}{2\pi} \int d\omega \,\omega^{2n} f^{W}(\omega)$$

and from them, the Lanczos coefficients b_n , according to

$$b_1^{2n} \cdots b_n^2 = \det(\mu_{i+j})_{0 \le i,j \le n}$$

where (μ_{i+i}) is a Hankel matrix of moments.

The *autocorrelation function* $C(t) := \varphi_0(t) \equiv \Pi^W(t, \mathbf{0})$ is the starting point for finding $\varphi_n(t)$, from which it is possible to compute the Krylov complexity $K_{\phi}(t)$.

The spectral density $ho(\omega, {f k})$ contains the physical information about the theory. For example, for a free massive scalar

The coefficients $a_n = (\tilde{\mathcal{O}}_n | \mathscr{L} | \tilde{\mathcal{O}}_n)$ and $b_n = (\tilde{\mathcal{O}}_{n-1} | \mathscr{L} | \tilde{\mathcal{O}}_n)$ $= (\tilde{\mathcal{O}}_n | \mathcal{L} | \tilde{\mathcal{O}}_{n-1})$ are called *Lanczos coefficients*.

If \mathcal{O} is Hermitian, then $a_n = 0$.

providing a tighter bound than the Maldacena-Shenker-Stanford (MSS) bound on Lyapunov exponents λ_L of OTOCs



$$\rho(\omega, \mathbf{k}) = \frac{\mathcal{N}}{\epsilon_{\mathbf{k}}} \left(\delta(\omega - \epsilon_{\mathbf{k}}) - \delta(\omega + \epsilon_{\mathbf{k}}) \right)$$

$$h \epsilon_{\mathbf{k}} := \sqrt{|\mathbf{k}|^2 + m^2}.$$

(4) Free massive scalar QFTs with <u>unbounded spectrum</u>

The power spectrum for a massive scalar field is given by

 $f^{W}(\omega) \sim \Theta(|\omega| - m)(\omega^2 - m^2)^{(d-3)/2} / |\sinh(\beta\omega/2)|$

The mass induces a dimerization, or "staggering", of the Lanczos coefficients b_n



(5) Free massive scalar QFTs with bounded spectrum

We can introduce a **UV cutoff** in the power spectrum, such that

 $f^{W}(\omega) = 0 \quad (|\omega| > \Lambda)$

This causes a **saturation** of the Lanczos coefficients



consistent with lattice computations ([Avdoskhin, Dymarsky & Smolkin (2022)]) and a transition in the Krylov complexity from an exponential to a **linear growth**:

(6) Interacting scalar in 4d

with

We consider two types of perturbations in d = 4:



a) Marginally irrelevant deformation ($\ell = 4$)

In this case, the one-loop self-energy is $\Pi_F =$ ____

This amounts to a change in the **effective mass** of the theory by a thermal mass contribution

$$m_{\text{eff}} = \sqrt{m^2 + m_{\text{therm}}^2} = \sqrt{m^2 + g_4/(24\beta^2)}$$

Thus, we obtain similar results to the free massive case (4).

b) Relevant deformation ($\ell = 3$) (massless case)

In this case, the one-loop self energy is $\Pi_E = ---($

The Lanczos coefficients exhibit a **staggering** that decreases with n $b_n(g)$

 $b_n \sim$ where: $|\gamma_{\text{odd}} - \gamma_{\text{even}}| \sim m$ $\frac{1}{\beta}n + \gamma_{\text{even}}$ (*n* even)

The presence of the mass breaks the smoothness of the Lanczos coefficients, and decreases the growth rate of the Krylov complexity:





The UV cutoff affects primarily the large-n behavior of the Lanczos coefficients b_n and the late-time behavior of Krylov complexity $K_{\phi}(t)$.



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